The Quark-Meson Coupling model as a description of dense matter

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Outline



- The Nature of Dense Matter
- The Models

2 Simulations:

- Hadronic Matter
- Mixed-Phase Matter



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The Nature of Dense Matter The Models

What We Know

- Quarks and Gluons are the fundamental degrees of freedom
- At low densities, Baryons (Nucleons) are the effective degrees of freedom
- At high densities...



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- At low densities, Baryons (Nucleons) are the effective degrees of freedom
- At high densities... ??? = Hyperons? Quarks? Other?



The Nature of Dense Matter The Models

Hadronic Models A Brief Overview

$\mathbf{Q} \text{uantum } \mathbf{H} \text{adro} \mathbf{D} \text{ynamics } (\mathbf{QHD}) \text{ Model}$

- Simple description of nucleons immersed in mean-field σ , ω , and ρ potentials,
- Constructed at the baryon level,
- Issues with large scalar potentials causing negative effective masses.

- Similar final form as QHD, but with self-consistent response to the σ field, despite construction from quark level,
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QMC: $M_B^* = M_B + \Sigma_B^s = M_B - w_B^s g_{\sigma N} \langle \sigma \rangle + \frac{d}{2} \tilde{w}_B^s \left(g_{\sigma N} \langle \sigma \rangle \right)^2$

The Nature of Dense Matter **The Models**

Hadronic Models A Brief Overview

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Hartree Σ_B^s (QHD)

$$\begin{split} \Sigma_B^s &= -g_{\sigma B} \langle \sigma \rangle \\ &= -g_{\sigma B} \sum_{B'} \frac{g_{\sigma B'}}{m_{\sigma}^2} \frac{(2J_{B'}+1)}{(2\pi)^3} M_{B'}^* \int \frac{\theta(k_{F_{B'}} - |\vec{k}|) d^3k}{\sqrt{\vec{k}^2 + M_{B'}^2}} \end{split}$$

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The Nature of Dense Matter The Models

Hyperonic QMC

- $B \in \{p, n, \Lambda, \Sigma^-, \Sigma^0, \Sigma^+, \Xi^-, \Xi^0\} = \{N, Y\}$
- $\ell \in \{e^-, \mu^-\}$
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Saturation

$$\begin{array}{ll} (E/A)_{\rho_0} &= -15.86 \text{ MeV}, \\ (\rho_{\text{total}})_{\rho_0} &= 0.16 \text{ fm}^{-3} \end{array} \right\} \quad -- \quad g_{\sigma N}, \ g_{\omega N} \\ (a_{sym})_{\rho_0} &= 32.5 \text{ Mev} \qquad -- \quad g_{\rho N} \end{array}$$



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Effective masses from Ref. [4]: Guichon *et. al.* doi:10.1016/j.nuclphysa.2008.10.001 (previously from Ref. [5]: Rikovska-Stone *et. al.* doi:10.1016/j.nuclphysa.2007.05.011) derived from the bag model.
The Nature of Dense Matter The Models

Hyperonic QMC

Equation of State (EOS) is calculated assuming that

$$\mu_i = B_i \, \mu_n - Q_i \, \mu_e = \sqrt{k_{F_i}^2 + (M_i + \Sigma_i^s)^2 + \Sigma_i^0}$$



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The Nature of Dense Matter The Models

Hadronic Models A Brief Overview: Finite Nuclei

Finite Nuclei:



The Nature of Dense Matter The Models

Hadronic Models A Brief Overview: Finite Nuclei

The mean-fields $\langle m \rangle$ are calculated via the equations of motion;

Equations of Motion

(C

$$= + m_{\sigma}^{2} \sigma = g_{N\sigma} \bar{\psi} \psi,$$

$$\frac{\partial^{\mu} \Omega_{\mu\nu}}{\partial^{\mu} R_{\mu\nu}^{a}} = g_{N\omega} \bar{\psi} \gamma_{\nu} \psi - m_{\omega}^{2} \omega_{\nu},$$

$$\frac{\partial^{\mu} R_{\mu\nu}^{a}}{\partial^{\mu} R_{\mu\nu}^{a}} = g_{\rho} \bar{\psi} \gamma_{\nu} \tau^{a} \psi - m_{\rho}^{2} \rho_{\nu}^{a}.$$



The Nature of Dense Matter **The Models**

Hadronic Models A Brief Overview: Finite Nuclei

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The Nature of Dense Matter The Models

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The mean-fields $\langle m \rangle$ are calculated via the equations of motion;

$$\left(-\nabla^2 + m_{\sigma}^2\right)\sigma(x) = -g_{N\sigma}\mathrm{Tr}[iG_H(x,x)],$$

$$\left[-\nabla^2 + m_{\omega}^2\right)\omega^{\mu}(x) = -g_{N\omega}\operatorname{Tr}[i\gamma^{\mu}G_H(x,x)]$$

$$\left(-\nabla^2 + m_{\rho}^2\right)\rho^{\mu a}(x) = -g_{\rho} \operatorname{Tr}[i\tau^a \gamma^{\mu} G_H(x,x)]$$



The Nature of Dense Matter **The Models**

Hadronic Models A Brief Overview: Finite Nuclei

Consider the solutions of the Dirac equation to be written as

Dirac Solutions

$$U_{\alpha}(x) = U_{n\kappa m t}(x) = \begin{bmatrix} iG_{n\kappa t}(r)/r & \Phi_{\kappa m}\eta_t \\ -F_{n\kappa t}(r)/r & \Phi_{-\kappa m}\eta_t \end{bmatrix}$$



The Nature of Dense Matter The Models

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Hadronic Models A Brief Overview: Finite Nuclei

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- ϵ_{α} , masses/splittings
- hypernuclei data
- \Rightarrow Compare to experiment!



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Hadronic Matter Mixed-Phase Matter

Hyperonic QMC (2007)



Hadronic Matter Mixed-Phase Matter

Hyperonic QMC (2008)



The improvement in the 2008 parameterization of M^* is that the effect of the mean scalar field in-medium on the familiar one-gluon-exchange hyperfine interaction (that in free space leads to the N- Δ and Σ - Λ mass splittings) is also included self-consistently.

This has the effect of increasing the splitting between the Λ and Σ masses as the density rises and the prime reason why we find that the Σ hypernuclei are unbound.

Guichon, Thomas, Tsushima: 2008



Hadronic Matter Mixed-Phase Matter

QMC - Finite Nuclei



[e.g. from C. E. Price, G. E. Walker: Phys.Rev.C36:354-364 (1987)]

Hadronic Matter Mixed-Phase Matter

QMC - Finite Nuclei



J. D. Carroll QMC dense matter

Hadronic Matter Mixed-Phase Matter

QMC - Finite Nuclei



J. D. Carroll QMC dense matter

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Hadronic Matter Mixed-Phase Matter

QMC - Finite Nuclei

	$\begin{array}{c} E_B \ (\mathrm{MeV}) \\ [\mathrm{experiment}] \end{array}$	$\begin{array}{c} E_B \ (\text{MeV}) \\ [\text{QMC}] \end{array}$	$r_c \text{ (fm)}$ [experiment]	$r_c (fm)$ [QMC]
$^{16}\mathrm{O}$	7.976	7.618	2.73	2.702
⁴⁰ Ca	8.551	8.213	3.485	3.415
⁴⁸ Ca	8.666	8.343	3.484	3.468
$^{208}\mathrm{Pb}$	7.867	7.515	5.5	5.42

[from H. H. Matevosyan: Ph.D. Thesis (2007)]



Hadronic Matter Mixed-Phase Matter

Tolman-Oppenheimer-Volkoff

Equations describe a static, spherically symmetric, non-rotating star, stable against gravitational collapse;



Hadronic Matter Mixed-Phase Matter

Tolman-Oppenheimer-Volkoff

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Hadronic Matter Mixed-Phase Matter

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$\underset{\rm TOV \ solutions}{\rm Hyperonic} \ QMC$



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J. D. Carroll QMC dense matter

Hadronic Matter Mixed-Phase Matter

What about higher densities?

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- 3 quarks in a 'bag',
- Separated from the QCD vacuum by an energy-density B,
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- Simple inclusion of $D\chi SB$,
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Quark Models A Brief Overview

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MIT:

$$m_u=5~{\rm MeV},~m_d=7~{\rm MeV},~m_s=95~{\rm MeV}$$

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NJL:

$$m_q^* = m_q + \Sigma_q^s = m_q - 2G\langle \bar{\psi}_q \psi_q \rangle$$

= $m_q + \frac{8G\mathcal{N}_c}{(2\pi)^3} \int \frac{\theta(k_F - |\vec{k}|)\theta(\Lambda - k_F)m_q^s}{\sqrt{\vec{k}^2 + m_q^{*2}}}$

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NJL: $k_F = 0$: $m_u = 350 \text{ MeV}, m_d = 350 \text{ MeV}, m_s = 450 \text{ MeV}$ $k_F = \Lambda$: $m_u = 5 \text{ MeV}, m_d = 7 \text{ MeV}, m_s = 95 \text{ MeV}$



Hadronic Matter Mixed-Phase Matter

Quark Models NJL Effective Masses



Hadronic Matter Mixed-Phase Matter

Phase Transitions



Hadronic Matter Mixed-Phase Matter

Phase Transitions

Gibbs Conditions	
• $T_H = T_Q$	– Thermal Equilibrium
• $(\mu_i)_H = (\mu_i)_Q$	– Chemical Equilibrium
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Hadronic Matter Mixed-Phase Matter

Hyperonic QMC Phase Transition



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The Basics: Simulations: Hadronic Matter Mixed-Phase Matter

$\begin{array}{c} Hyperonic \ QMC \\ Phase \ Transition \end{array}$



Hadronic Matter Mixed-Phase Matter

Quark Chemical Potentials related to independent chemical potentials;

Chemical Equilibrium - quarks

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$\begin{array}{c} Hyperonic \ QMC \\ Phase \ Transition \end{array}$



Hadronic Matter Mixed-Phase Matter

Nucleonic QMC Phase Transition



Hadronic Matter Mixed-Phase Matter

Mixed-Phase Hyperonic QMC TOV solutions



- The inclusion of $D\chi SB$ prevents a phase transition to quark matter,
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Hadronic Models A Brief Overview: Hartree–Fock

At Hartree–Fock level, the scalar self-energy also includes an exchange term, and becomes momentum-dependent:





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Hartree–Fock $\Sigma_B^s(k)$ (QHD)

$$\begin{split} \Sigma_B^s(k) &= -g_{\sigma B} \langle \sigma \rangle \\ &+ \frac{1}{4\pi^2 k} \int_0^{k_{F_{B'}}} \frac{q \ M_B^*(q)}{E_B^*(q)} \\ &\times \left[\frac{1}{4} g_{\sigma B'}^2 \Theta_{\sigma}(k,q) - g_{\omega B'}^2 \Theta_{\omega}(k,q) \right] \ dq \end{split}$$



3MI

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3WI

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• Momentum-dependent $\Sigma \Rightarrow$ harder to solve

- re-definition of μ_i
- changes to self-consistencies



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Hadronic Models A Brief Overview: Hartree–Fock

OR... $m = \langle m \rangle + \frac{\delta m}{\delta m}$



Hadronic Models A Brief Overview: Hartree–Fock

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Hadronic Models A Brief Overview: Hartree–Fock

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Hartree–Fock
$$H_s$$
 (QHD)
$$H_{\sigma} = \int d\vec{r} \left[E(\langle \sigma \rangle) - \frac{1}{2} \langle \sigma \rangle \left\langle \frac{\partial E}{\partial \langle \sigma \rangle} \right\rangle + \frac{1}{2} \delta \sigma \left(\frac{\partial E}{\partial \langle \sigma \rangle} - \left\langle \frac{\partial E}{\partial \langle \sigma \rangle} \right\rangle \right) \right]$$



Hadronic Models A Brief Overview: Hartree–Fock

- still momentum-independent (additional energy contribution)
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Further Reading I

🍆 Carroll

Applications of the Octet Baryon Quark-Meson Coupling Model to Hybrid Stars (PhD Thesis). arXiv:1001.4318

Carroll, Thomas The Hyperfine, Hyperonic QMC Model - Extension to Hartree–Fock I: Infinite Nuclear Matter. in preparation

 Carroll, Leinweber, Williams, Thomas Phase Transition from QMC Hyperonic Matter to Deconfined Quark Matter. Phys.Rev.C79:045810, 2009 [doi:10.1103/PhysRevC.79.045810]



Further Reading II

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 Nucl.Phys.A814:66-73, 2008
 [doi:10.1016/j.nuclphysa.2008.10.001]
- Rikovska-Stone, Guichon, Matevosyan, Thomas Cold uniform matter and neutron stars in the quark-mesons-coupling model. Nucl.Phys.A792:341-369, 2007
 [doi:10.1016/j.nuclphysa.2007.05.011]



	Appendix	Further Reading Thank You
Thank You!		

 $\mathcal{F}\mathit{in}$

